

# History of Modern Astronomy

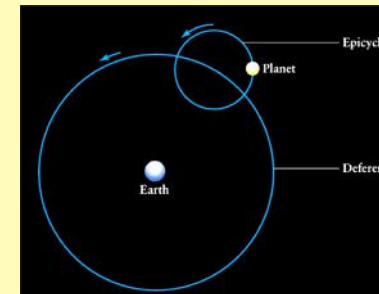


## Ptolemy (100 -170 ACE)

- *synthesized* others' work into a *comprehensive model*
- included *Sun & Moon, Mercury, Venus, Earth, Mars, Jupiter & Saturn*
- *geocentric (Earth centered) model of the solar system*



- used for *1,500 years even though it did not correctly predict planetary motions!*
- planets typically move *eastward* across the sky
- *periodically they appear to stop, move westward for a while, then stop and move eastward again (eg) Mars, 2003*
- *westward motion is called retrograde motion*



- *Ptolemy added epicycles to his model to attempt to account for retrograde motion*

- need *different epicycles, periods for each planet*
- complex! "*Occam's Razor*" (14th century Franciscan friar)

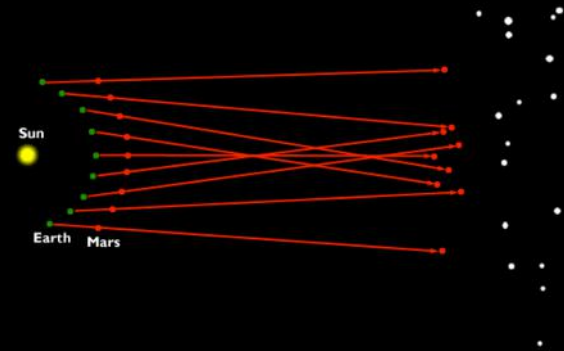


## Nicolaus Copernicus (1473-1543)



- Ptolemy's *geocentric model* was *complex & inaccurate*
- revived *Greek idea of heliocentric (Sun centered) model* of the solar system in *1543*
- offered *simple* explanation for *retrograde motion* but *didn't* improve planetary motions otherwise

## Retrograde explained



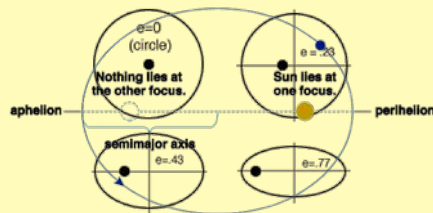
## Johannes Kepler (1571 - 1630)



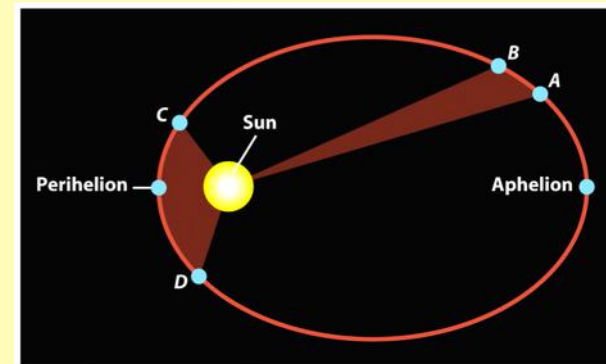
- took *13 years* to develop *3 laws of motion* for *heliocentric model*

### Kepler's Laws of Planetary Motion

1. planets move around the Sun in *elliptical* orbits

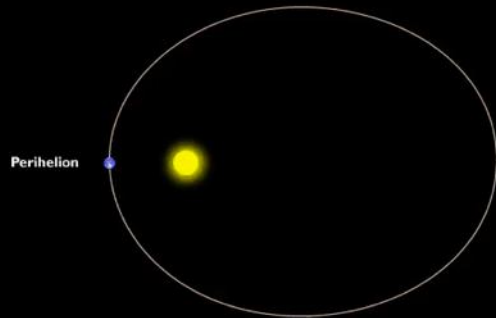


2. A line joining a planet and the Sun will "sweep out" *equal areas in equal amounts of time*



- planets move *faster close to the Sun (perihelion)*

## Kepler's 2nd Law



3. The *time* for a planet to orbit the Sun (*period, P*) depends on its *average distance from the Sun (a)*:

$$P^2 = a^3$$

- *period* measured in *years* & *semi-major axis (a)* measured in *astronomical units (AU)*

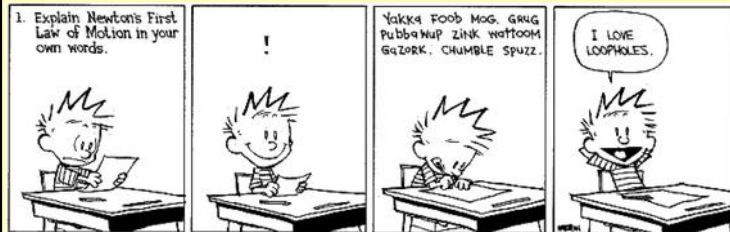
(eg) Jupiter:  $a = 5$ ,  $P = \sqrt{(5)^3} = \sqrt{125} \approx 11$  years

**CLICKER:** For Saturn,  $a \approx 10$ , so  $P$  is *roughly*  
(a) 5 years (b) 10 years (c) 30 years (d) 1000 years

## Isaac Newton

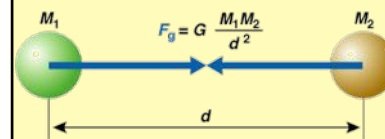
(1642 - 1727)

- *invented reflecting telescope*
- *invented calculus*
- *deduced gravitation & basic laws of motion*

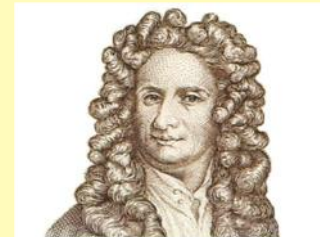


## Law of Gravitation

- *any objects* with mass *attract each other*



- *attraction occurs even over astronomical distances*



**Gravity.**

It's not just a good idea.  
It's the Law.

$F_g = G \frac{M_1 M_2}{d^2}$

**CLICKER:** What is the effect on the force of gravity if two objects are moved twice as far apart?

- (a) the force is one quarter as big
- (b) the force is cut in half
- (c) the force is doubled
- (d) no impact (gravity is a constant)

- whirl a ball at the end of a string...
- ball *wants* to move in *straight line*
- keep it moving in a circle by *exerting a force*

- Sun's **gravity keeps planets in orbit**
- Earth does the same thing to satellites, etc.

**Q:** Why don't things in orbit just fall to Earth?

**DEMO:** drop an object vs throw an object

**A:** They do! But they also move "sideways" fast enough to avoid getting closer to the Earth.

**DEMO:** a bucket of water

Not to scale!

