

Basic informationDensity (avg): 3.9 g/cm³Distance from Sun (avg): 1.52 AUOrbital Period: 687 Earth daysRotation Period: 24 hrs, 36 minAlbedo: 0.25Moons: 2Atmosphere: much thinner than Earth'sTilt of rotational axis: 24°

Mars "The Red Planet"

- named for *Roman god* of *war*
- *reddish* color from *iron-oxide* ("rust")
- *visible* surface features: polar ice caps, volcanoes, changing dark regions



Exploring Mars

• Mars (1960's-70's) – mostly failed USSR • Mariner 4&9 (1965,1971) & Viking (1976) NASA • resolutions from 1 km to 100 m (10 m) • Mars Global Surveyor ; 2-10 m (1997) NASA NASA • MER Spirit, Opportunity (2003) • Mars Express (2 m in stereo) (2003) ESA • **MRO** (0.25 m) (2006) NASA • **MSL** "Curiosity" (2012) NASA • **MOM**, **MAVEN** (2014) ISRO, NASA • **TGO** (2016) ESA/USSR • InSight (2018) NASA



Mars (Viking, 1976-82)







Martian Canals

• *Huygens* sketched dark region *Syrtis Major* (1659)



• *Giovanni Schiaparelli* noticed a network of dark lines on surface, named them "*canali*" (*channels*)

- Percival Lowell: "Martian Canals" (late 1800's)
- patterns of light & dark material
- *dust storms* blow *light "sands"* around, exposing *darker rock*



Dust Storms





Interior Structure

- *little seismic data* (inSight!)
- smalle size, lower density
- smaller core, "solid" interior?
- remnant magnetic field
- past evidence of plate tectonics?
- Valles Marineris: largest & deepest canyon system in solar system could be a plate boundary









• *Olympus Mons:* largest volcano in the solar system (27 km high or 3x height of Mt. Everest)





Q: Why is it so large? (would cover 1/4 of BC!)



• low gravity, no active plate tectonics

Atmosphere

• pressure: 0.007 bars (<1% of Earth)

	Venus	Earth	Mars
Nitrogen (N ₂)	3.5%	78.08%	2.7%
Oxygen (O ₂)	almost zero	20.95%	almost zero
Carbon dioxide (CO ₂)	96.5%	0.035%	95.3%
Water vapor (H ₂ O)	0.003%	about 1%	0.03%
Other gases	almost zero	almost zero	2%

- trace O, CH_4
- source: volcanic... (biologic??)
- atmosphere *escaped into space (low gravity)*

Temperature • "Daytime" temperature: 240 K (-33°C) • "Nighttime" temperature: 170 K (-103° C) • ground temperature can exceed 0° C







- visual evidence
- *chemical evidence* from rovers (*hematite*, *sulfates*)
- salty, shallow seas (?) covered *northern lowlands*
- volcanic activity
- greenhouse effect
- \Rightarrow thicker, warmer atmosphere
- *Q*: *Where is the water now?*
- vaporized, *ice caps*, subsurface (*permafrost*)











Q: Why is *quality* on left *so bad*?

Martian "Cities" *Cydonia (Candor City*)





CLICKER: Evidence that Mars once possessed significant water on its surface is ...?

(a) minerals found in rocks examined by rovers
(b) significant water ice found in Mars' ice caps
(c) ancient flow patterns on the surface
(d) all of the above