

Basic information

Average Density: 1.4 g/cm³

Luminosity (J/s): a *trillion-trillion* 100 W bulbs!

Rotation Period (equatorial): 25 days

Surface temperature: ~5800 K

Core temperature: ~15,000,000 K

The Sun (Sol)

• **NEVER** look directly at the Sun (*eg. Galileo*)

• an *average* star & the *only one* in our solar system

• composed of *hot gas* (mostly *H*, *He*)

• contains 99.9% of the mass of *solar system*

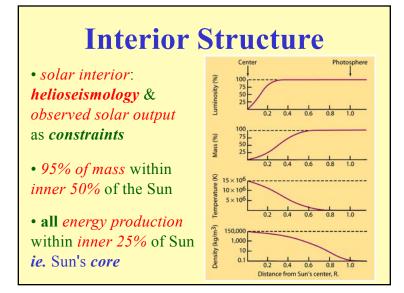
Q: Do planetary orbits make sense given this?

Size

- 109 times the diameter of Earth
- about 1 million Earths could fit inside Sun

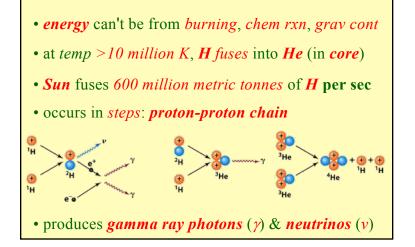


• ~330,000 times as massive as Earth

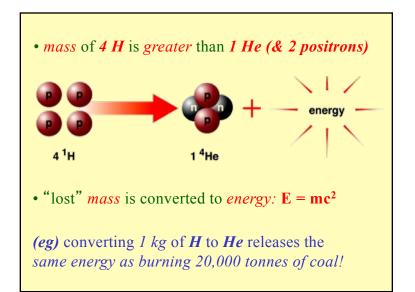


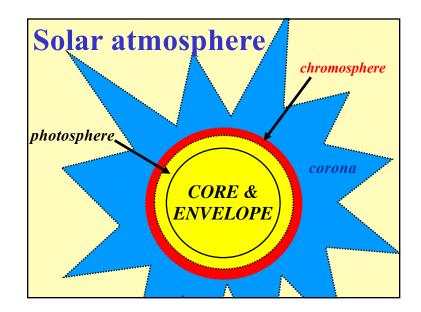


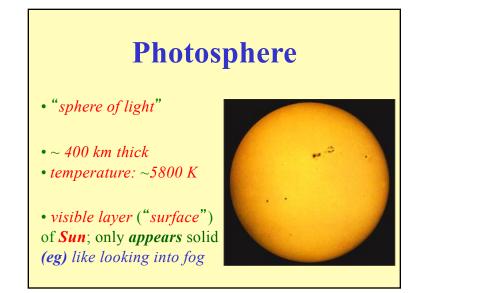
only radiation & convection operate inside Sun
mode depends on temperature, density, gradient
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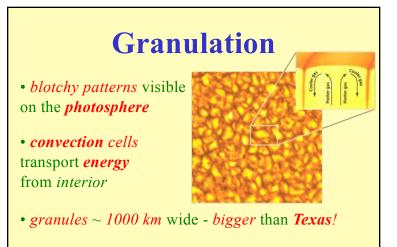


Fusion

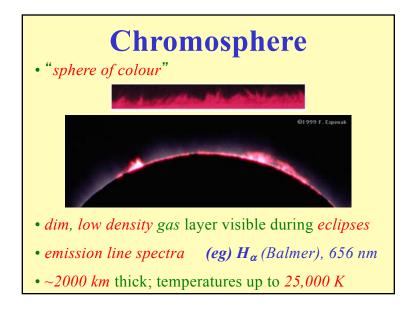


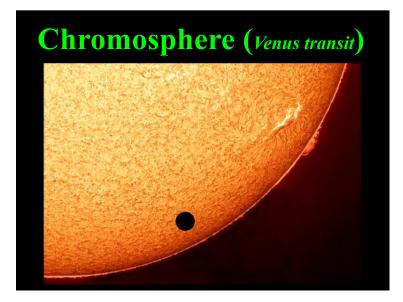




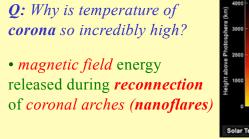


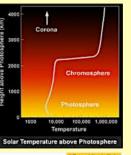
Q: Why are they dark *along their edges*?



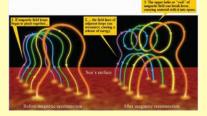


Corona • outermost layer • very high temperature, very low density gas • emission line spectra • ionized species ⇒ high temperatures • 1930's: 1-2 million K Q: High temperature but not "hot" ... ???









• corona emits UV & X-rays

only one millionth as bright as *photosphere*"like" *full moon*

CLICKER: The corona is very dim relative to the photosphere because...?

(a) most of the emission is not in visible light

(b) the density is very low

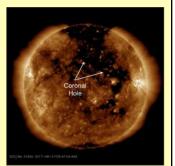
(c) it contains colder gas than the photosphere

(d) magnetic field lines are unable to trap hot gas

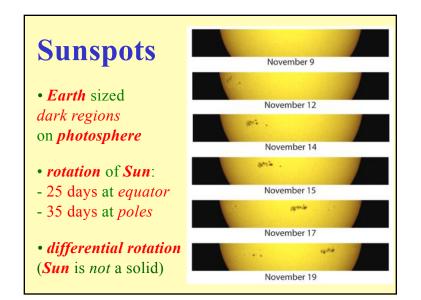
(e) both a & b (f) both b & c (g) both c & d

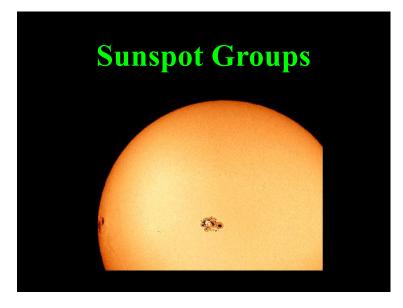
The Solar Wind

- mass loss from corona
- electrons, H & He nuclei
- one million tonnes/second
- *preferential loss* via open field lines: *coronal holes*
- average speed & density: 450 km/s & < 5 per cm³

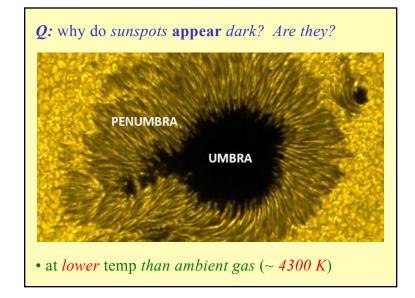


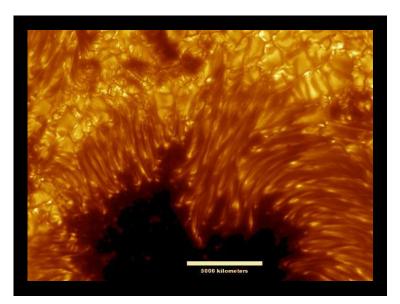
• SW loss ~ 0.1 % of **Sun's** mass over lifetime





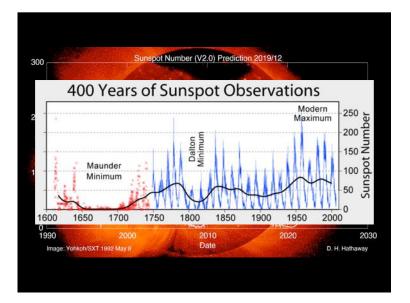
Sunspots & granulation

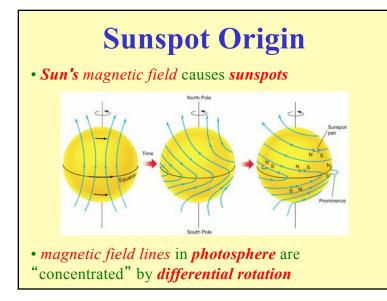


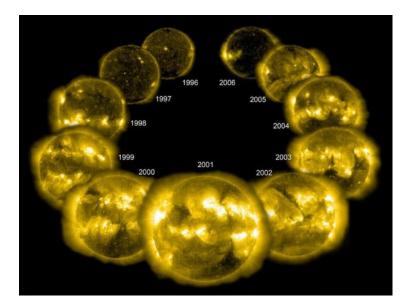


Solar Cycle

- Sun has a magnetic field of varying strength (eg) Earth:0.5 G; Sun:1-100 G; sunspots:1000's G
- sunspot formation related to magnetic field
- sunspots: regions of concentrated magnetic field
- sunspot cycle takes ~11 years, max to max
- magnetic field orientation: 22 year solar cycle





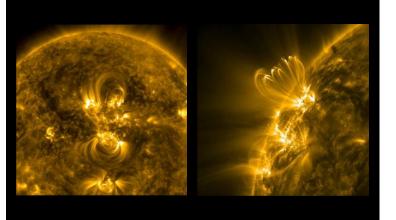


- convection causes field lines to tangle & twist
 kinks form, poking through photosphere
 kinked field line = intense magnetic field
 sunspots form at base of each loop (in pairs)
 hot plasma repelled by magnetic field ⇒ cooler
 - **Q:** Why do **sunspots** eventually disappear?

- *prominences*: loops of gas from chromosphere
 eventually fall back into Sun
- *Q*: Why are they *loops*?
- *filaments*: *prominences* viewed from *above*
- *plages*: bright regions in *chromosphere* prior to appearance of *sunspots*

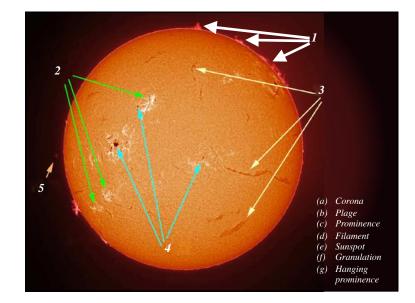






Hanging Prominence





Flares & CMEs

- *speed & density*: up to 1000 km/s, 100 cm⁻³
- may penetrate Earth's magnetic field & cause electrical disruptions (eg) 1989 Hydro Quebec
- classes: A<B<C<M<X
- associated with sunspots

