Neutron Stars: The Basics and Beyond



Figure 1: An artist's conception of a binary neutron star system. Retrieved from https://www.psu.edu/news/ research/story/seeing-light -neutron-star-collisions

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Table of Contents

1. About Neutron Stars

- a. History
- b. The fate of a star
- c. Formation
- d. Properties
- e. Classifications
- 2. A Hypothetical Use of Neutron Stars
 - a. Analyzing the Properties of Neutron Stars
 - b. What If We Built a Computer out of Neutron Stars?
- 3. Another Way to Use Neutron Stars
 - a. Some Dustin idea involving the frictionless nature of NS
 - b. Why These Ideas Are Totally Infeasible
- 4. Bookkeeping
 - a. References
 - b. Q&A

First detected in 1967 by Jocelyn Bell and Antony Hewish (PBS, n.d).

Initially thought to be signals from extraterrestrial civilization, dubbed "Little Green Men" (Menezes, 2021).

By 1968, identified as a neutron star and later named Cambridge Pulsar (PBS, n.d.).

History



Figure 2: Bell and Hewish. Retrieved from https://doi.org/10.1038/547005b

A star's fate depends on its mass. It may become a white dwarf, a neutron star, or a black hole (Goddard Space Flight Center, 2015).

High-mass stars, with initial mass greater than 8 times the Sun, evolve into red supergiants, which undergo a supernova event (European Space Agency, n.d).

The Fate of a Star

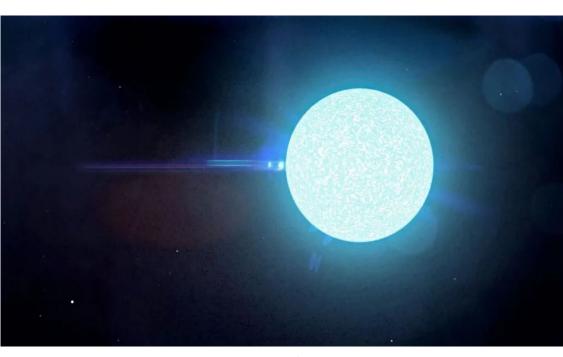


Figure 3: An artist's conception of a neutron star. Retrieved from https://www.sciencenews.org/article/neutron-stars-pulsar-

https://www.sciencenews.org/article/neutron-stars-pulsardensity-diameter-matter

The Fate of a Star, Cont.



Figure 4: An artist's conception of a "New Black Neutron Star". Retrieved from https://globalnews.ca/news/7102112/black-neutron-star-ligo-hole/ If the core that is left is between 1 and 3 solar masses, it forms a neutron star; otherwise, it becomes a black hole (European Space Agency, n.d).

Formation of a Proto-Neutron Star

Strong Nuclear Force: Binds protons and neutrons and holds the nucleus together (Nobel Prize Outreach, 2025).

Weak Nuclear Force: Enables proton-neutron conversions via radioactive decay (Sears & Zemanski, 2010, as cited in Donev et al., 2014).

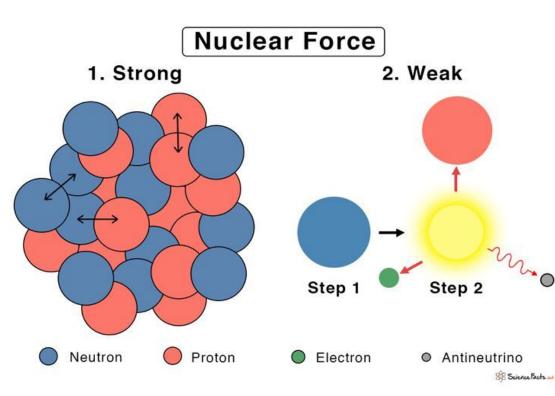


Figure 5: Illustration of the strong and weak forces. Retrieved from https://www.sciencefacts.net/nuclear-force.html

Formation of a Proto-Neutron Star, cont.

Timeline

Within 10 seconds of core collapse: PNS forms, neutrino emissions, and rapid contraction. (Camelio, 2017)

After 10 seconds:

Becomes quasi-stationary with strong thermodynamic gradients and neutrino emissions (Martinion et al., 2014).

Contracts and cools

Forms into a full neutron star.

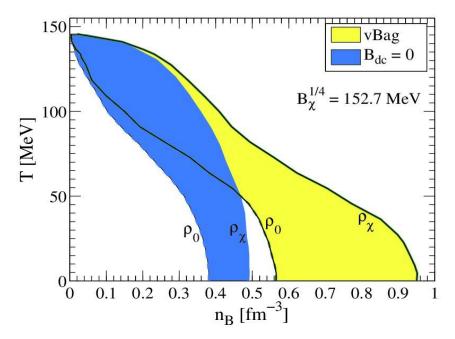


Figure 6: Illustration of a density vs. temperature graph for proto-neutron states. Retrieved from https://doi.org/10.1088/1742-6596/861/1/012026

Physical Qualities of a Neutron Star

Density

Neutron stars are extremely compact, spanning just 10 – 20 km in diameter (Alfredo, 2023). They typically have a mass 1.1 to 2.3 times that of the Sun.

Temperature

The temperature of a proto-neutron star is measured to be between 10¹¹ – 10¹² Kelvin. Because of the neutrinos emissions, the temperature falls and settles to about 10⁶ Kelvin (Camelio, 2017) for a neutron star.

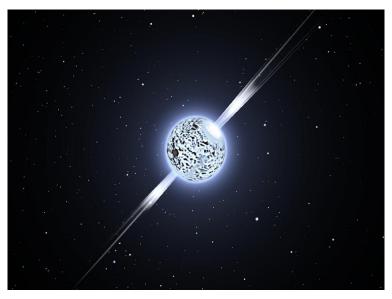


Figure 7: An artist's conception of a neutron star. Retrieved from http://www.orionsarm.com/eg-article/46d4bf6c4e652

Physical Qualities of a Neutron Star, Cont.

Magnetic Fields

Neutron stars have magnetic fields ranging from 10⁸ to 10¹⁵ G (Reisenegger, 2001). These intense magnetic fields can emit bursts of radiation, influencing their surrounding environments (Vigano, 2013, as cited in Alfredo, 2023).

Rotation

Neutron stars can rotate up to hundreds of times per second due to angular momentum conversion during the core collapse of the previous star (Ordaz, 2019, as cited in Alfredo, 2023).

Interior Composition

Neutron stars contain superfluids and superconductors in their internal environment (Page et al., 2014, as cited in Chamel, 2017)

Main types: (Jet Propulsion Laboratory, 2020)

- I. Magnetars
- II. Pulsars
- III. Isolated Neutron Stars
- **Pulsar Subtypes:** (Center for Astrophysics: Harvard & Smithsonian, n.d.)
- I. Millisecond
- II. Ordinary

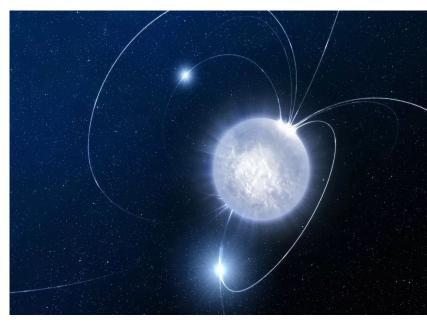


Figure 8: An artist's conception of neutron star. Retrieved from https://rndmcanada.org/2023/12/11/neutron-star/

Classifications

Clicker Question

What force is primarily responsible for holding a neutron star together?

A): Electromagnetic force

B): Weak nuclear force

C): Strong nuclear force

D): Gravitational force

Analyzing the Properties of Neutron Stars



Figure 9: An artist's conception of a neutron star. Retrieved from https://www.cnet.com/science/this-astounding-neutron-s tar-spins-faster-than-a-blender/ Superconductivity allows for:

- Phones that never die (U.S Department of Energy, n.d.).
- Computers that never overheat (U.S Department of Energy, n.d.).

Superconductivity is a highly sought-after property for a material, but we've only achieved it at very low temperatures (Wu et al., 1987).

Well, Neutron Stars are superconductive, and certainly not at low temperatures...

What if We Built a Computer Out of Neutron Stars?

- iPhone 15 Pro Max's chip has 19 billion transistors (CNET, n.d.)

 No sources on transistor volume, so let's say 1 cubic micrometer to keep things simple.

 1 cubic micrometer times 19 billion gives us 0.019 cubic centimeters of materials we need to supplant with Neutron Star stuff. A single M&M is 33 times the size (Wolfram Alpha, n.d.)

- Now we have an electronic device that will never die!

Clicker Question

- How Might You Realistically Befriend a Worm?
- A) Be a good conversationalist.
- B) BYOB!
- C) Forage huckleberries and blackberries from the mountain as a gift
- D) err::MicrosoftInjectionError! [Please log out of the program.]

Other Things We Could Do with Neutron Stars

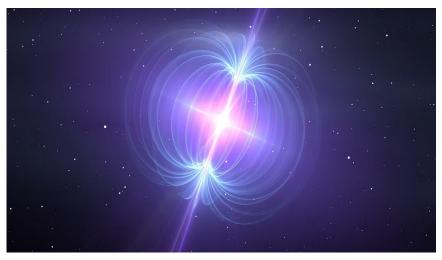


Figure 10: An artist's conception of a neutron star. Retrieved from http://theconversation.com/this-newly-discoveredneutron-star-might-light-the-way-for-a-whole-n ew-class-of-stellar-object-184050 Frictionless nature of neutron star material;

Frictionless high speed modes of travel?

Airliners, high-speed rail?

Possibly frictionless airplanes?

Neutron Star Applications

Concorde reborn? But faster and more efficient?

Maybe not burning 25,000 L of fuel per hour? Mother Nature rejoice!



Figure 11: The Concorde. Retrieved from https://www.popularmechanics.com/flight/airline s/a27206102/concorde-badass-plane/

Why These Ideas Are Totally Infeasible

The nearest Neutron Star is 130 parsecs away, or 424 light years (Posselt et al., 2008).

Although we'd need only 1/33rd of an M&M's worth of material for an iPhone, it would weigh as much as the Great Pyramid of Giza (Wolfram Alpha, n.d.)

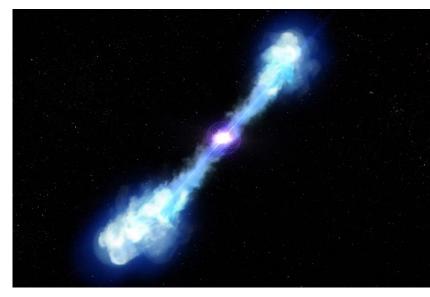


Figure 12: An artist's conception of gamma ray bursts from neutron stars. Retrieved from https://newatlas.com/space/neutron-star-collision-short-g amma-ray-burst/

Why It Is Not Happening, Cont.

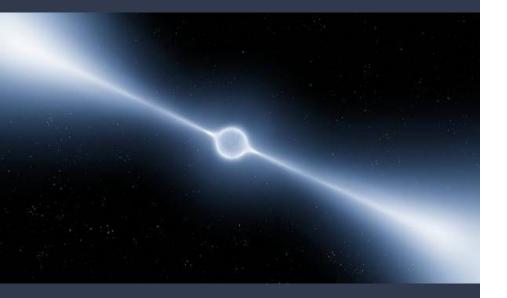


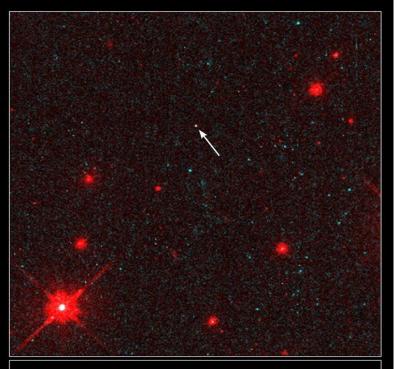
Figure 13: Neutron star stock photo. Retrieved from https://www.istockphoto.com/photos/neutron-star

Also, Neutron Stars only have these qualities (superfluidity, superconductivity) due to the immense pressure and temperature of their cores (Lombardo & Schulze, 2000).

Image: Neutron Star Trajectory



Figure 14: Photo of a neutron star trajectory in visible light. Retrieved from https://hubblesite.org/contents/news-release s/2000/news-2000-35.html Figure 15: A lone neutron star in space. About 28 km across. Retrieved from https://esahubble.org/images/opo9732a/



Isolated Neutron Star RX J185635-3754 Hubble Space Telescope • WFPC2

PRC97-32 • ST ScI OPO • September 24, 1997 F. Walter (State University of New York at Stony Brook) and NASA

Image: Isolated Neutron Star

Clicker Question

- What property of neutron star material would be beneficial for aircraft fuel efficiency?
- A): Super Conduciveness
- B): Superfluidity
- C): Superconductivity
- D): Super super supper

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Image References

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Question Period

Please ask better questions than they do in the real Question Period in Parliament.

I beg of you